

1 Part I: do 3 of 4

1. Suppose that $\rho(\mathbf{x}, t)$ and $\mathbf{v}(\mathbf{x}, t)$ solve the compressible Euler system

$$\begin{cases} \frac{\partial \mathbf{v}}{\partial t} + (\nabla \mathbf{v})\mathbf{v} = -\frac{\nabla(p(\rho))}{\rho}, \\ \frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = 0, \end{cases}$$

where $p(\rho) = \rho^{7/5}$. Assume that the flow is irrotational i.e. $\mathbf{v}(\mathbf{x}, t) = \nabla \phi(\mathbf{x}, t)$ for some scalar function ϕ . For what function $Q(\rho)$ is the quantity

$$E(\mathbf{x}, t) = Q(\rho(\mathbf{x}, t)) + \frac{\partial \phi}{\partial t}(\mathbf{x}, t) + \frac{1}{2} |\nabla \phi(\mathbf{x}, t)|^2$$

independent of \mathbf{x} ? Prove your assertion.

2. The period T of revolution of a planet moving along a circular orbit around a star depends only on the planet's distance to the star R and the star's mass M . Use dimensional analysis to find Kepler's formula for the ratio of the periods T_1/T_2 for any two such planets revolving around the same star.
3. Use scaling to express the wave function $\psi(\mathbf{x}, t)$, solving Schrödinger's equation

$$i\hbar \psi_t = -\frac{\hbar^2}{2m} \Delta \psi - \frac{e^2}{4\pi\epsilon_0 |\mathbf{x}|} \psi, \quad \mathbf{x} \in \mathbb{R}^3, t > 0$$

in terms of the solution $\phi(\boldsymbol{\xi}, \tau)$ of

$$i\phi_\tau = -\Delta \phi - \frac{\phi}{|\boldsymbol{\xi}|}, \quad \boldsymbol{\xi} \in \mathbb{R}^3, \tau > 0.$$

Give expressions for the characteristic time and length in terms of the positive constants \hbar , m , e and ϵ_0 .

4. Assume that the electric and magnetic fields $\mathbf{E}(\mathbf{x}, t)$, $\mathbf{B}(\mathbf{x}, t)$, $\mathbf{x} \in \mathbb{R}^3$, $t > 0$ are given by

$$\mathbf{E}(\mathbf{x}, t) = \begin{cases} \mathbf{E}^+, & \mathbf{x} \cdot \mathbf{n} > 0, t > 0 \\ \mathbf{E}^-, & \mathbf{x} \cdot \mathbf{n} < 0, t > 0 \end{cases}, \quad \mathbf{B}(\mathbf{x}, t) = \begin{cases} \mathbf{B}^+, & \mathbf{x} \cdot \mathbf{n} > 0, t > 0 \\ \mathbf{B}^-, & \mathbf{x} \cdot \mathbf{n} < 0, t > 0 \end{cases},$$

where \mathbf{n} is the unit normal to the plane $\mathbf{x} \cdot \mathbf{n} = 0$ separating two media. Suppose that the vectors \mathbf{E}^\pm , \mathbf{B}^\pm are constant. What extra conditions do these four vectors have to satisfy in order for the following two Maxwell's equations

$$\begin{cases} \nabla \times \mathbf{E} = -\frac{1}{c} \frac{\partial \mathbf{B}}{\partial t}, \\ \nabla \cdot \mathbf{B} = 0 \end{cases}$$

to hold in $\mathbb{R}^3 \times (0, +\infty)$ in the sense of distributions?

2 Part II: do 2 of 3

1. Suppose that a point mass m in a plane is attached by two identical springs, each with spring constant k , to two given points A and B in the plane. Assume that all springs are at their natural length when the point mass is at the midpoint of AB . Also assume that there is no gravity and no friction.
 - (a) Identify degrees of freedom.
 - (b) Write the system's Lagrangian.
 - (c) Write the Euler-Lagrange equations of motion.
2. Compute the Green's function for the boundary value problem

$$\begin{cases} x^2 u''(x) + xu'(x) + u(x) = f(x), & x \in [1, e^\pi] \\ u(1) = 0, \\ u'(e^\pi) = u(e^\pi). \end{cases}$$

Write the solution $u(x)$ in terms of this Green's function.

3. Use Laplace's method to find the first two terms of the asymptotic expansion of

$$I(x) = \int_0^1 t^x \sin^2 t dt,$$

as $x \rightarrow +\infty$.